

# Teaching Notes: Thermodynamic Behavior of an Ideal Fermi Gas

## Overview

We are studying a gas of **non-interacting fermions** (particles with half-integer spin like electrons, protons, etc.). The main distinguishing feature of a Fermi gas is that **no two fermions can occupy the same quantum state**, due to the **Pauli exclusion principle**. This affects how the gas behaves at both low and high temperatures.

## 1. What is an Ideal Fermi Gas?

- A collection of identical, non-interacting fermions.
- Each fermion has mass  $m$ , enclosed in volume  $V$  at temperature  $T$ .
- Fermions obey **Fermi-Dirac statistics**.

## 2. Grand Canonical Partition Function (Q)

In the grand canonical ensemble:

$$\frac{PV}{kT} = \ln Q = \sum_i \ln(1 + ze^{-\beta\epsilon_i})$$
$$N = \sum_i \frac{1}{z^{-1}e^{\beta\epsilon_i} + 1}$$

where  $\beta = 1/kT$  and  $z = e^{\beta\mu}$ .

## 3. Continuum Approximation (Non-Relativistic)

Using density of states in 3D:

$$\frac{P}{kT} = \frac{g}{\lambda^3} f_{5/2}(z),$$
$$\frac{N}{V} = \frac{g}{\lambda^3} f_{3/2}(z)$$

where

- $g$  = internal degrees of freedom
- $\lambda = \left(\frac{h^2}{2\pi m kT}\right)^{1/2}$
- $f_n(z) = \frac{1}{\Gamma(n)} \int_0^\infty \frac{x^{n-1}}{z^{-1}e^x + 1} dx$

## 4. Equation of State and Energy

Eliminating  $z$ :

$$PV = \frac{2}{3}U$$
$$U = \frac{3}{2}NkT \cdot \frac{f_{5/2}(z)}{f_{3/2}(z)}$$

## 5. Specific Heat and Entropy

$$C_V = Nk \left[ \frac{15}{4} \frac{f_{5/2}(z)}{f_{3/2}(z)} - \frac{9}{4} \frac{f_{3/2}(z)}{f_{1/2}(z)} \right]$$
$$S = Nk \left[ \frac{5}{2} \frac{f_{5/2}(z)}{f_{3/2}(z)} - \ln z \right]$$

## 6. Classical Limit (High T / Low n)

When  $z \ll 1$ ,  $f_n(z) \approx z$ :

$$P = \frac{NkT}{V}, \quad U = \frac{3}{2}NkT, \quad C_V = \frac{3}{2}Nk$$

## 7. Degenerate Limit ( $T \rightarrow 0$ )

At  $T = 0$ :

- Occupation number:

$$\bar{n}(\varepsilon) = \begin{cases} 1 & \varepsilon < \varepsilon_F \\ 0 & \varepsilon > \varepsilon_F \end{cases}$$

- Fermi energy:

$$\varepsilon_F = \frac{\hbar^2}{2m} (3\pi^2 n)^{2/3}$$

- Internal energy:

$$U = \frac{3}{5}N\varepsilon_F$$

- Pressure:

$$P = \frac{2}{5}n\varepsilon_F$$

## 8. Low Temperature Expansion (Sommerfeld)

$$\mu(T) \approx \varepsilon_F \left( 1 - \frac{\pi^2}{12} \left( \frac{kT}{\varepsilon_F} \right)^2 \right)$$
$$C_V = \frac{\pi^2}{2} Nk \frac{T}{T_F}$$
$$S = \frac{\pi^2}{2} Nk \frac{T}{T_F} \rightarrow 0 \text{ as } T \rightarrow 0$$

## Summary

- Fermi statistics leads to completely different low-T behavior than classical gases.
- At  $T = 0$ , pressure and energy are still non-zero due to the exclusion principle.
- $C_V \propto T$  at low T, consistent with experiments.